
Differential Neutrino Emissivities and Rates in Astrophysical Systems

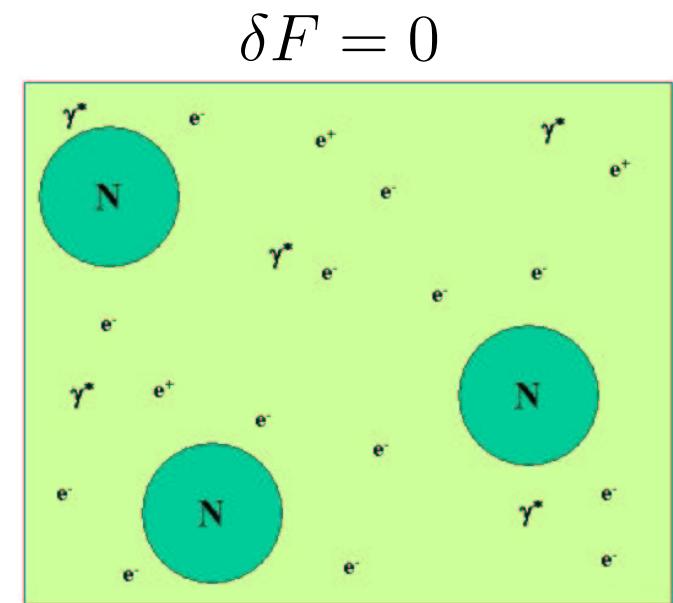
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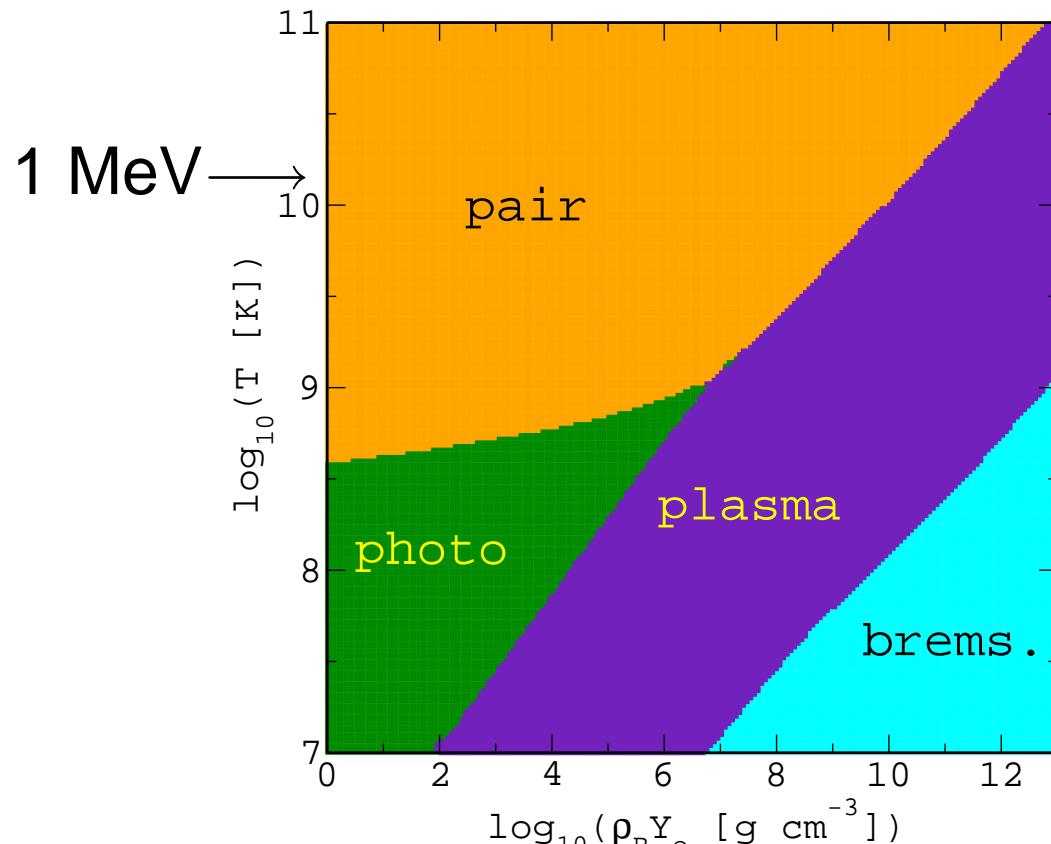
Physical setting

- ▶ neutrino emission
- ▶ Astrophysical systems
 - Degenerate He cores of red giant stars
 - Cooling in pre-white dwarf interiors
 - Accretion disks of GRB
 - Type II supernovae
 - Pulsar velocities
- ▶ QED plasma, subnuclear densities



Neutrino emitting processes at subnuclear densities

- ▶ pair neutrino
 $e^+e^- \rightarrow \nu\bar{\nu}$
- ▶ photo-neutrino
 $e^\pm\gamma \rightarrow e^\pm\nu\bar{\nu}$
- ▶ plasma neutrino
 $\gamma^* \rightarrow \nu\bar{\nu}$
- ▶ nucleon-nucleon
bremsstrahlung
 $NN \rightarrow NN\nu\bar{\nu}$



$$n_0 = 2.65 \cdot 10^{14} \text{ g cm}^{-3}$$

- Relative importance depends on s , T , ρ_B , and Y_e

Motivation and objectives

- ▶ Differential rates ($d\Gamma$) and emissivities (dQ) – crucial for ν transport a la Boltzmann equation
- ▶ In previous works Lenard's identity:

$$\int \frac{d^3 q}{2q_0} \frac{d^3 q'}{2q'_0} \delta^4(K - Q - Q') Q'_\mu Q_\nu = \frac{\pi}{24} \Theta(K^2) (2K_\mu K_\nu + K^2 g_{\mu\nu})$$

- Eliminates information about E & θ dependence of ν 's
- Bypasses calculation of $\langle |\mathcal{M}|^2 \rangle$
- ▶ Obtain $\langle |\mathcal{M}|^2 \rangle$, $d\Gamma$, dQ , and Legendre coefficients for plasma & photo-neutrino processes
- ▶ Checks for Q from previous works

Boltzmann transport equation

- Evolution of the ν distribution function $f(\mu_1, E_1)$

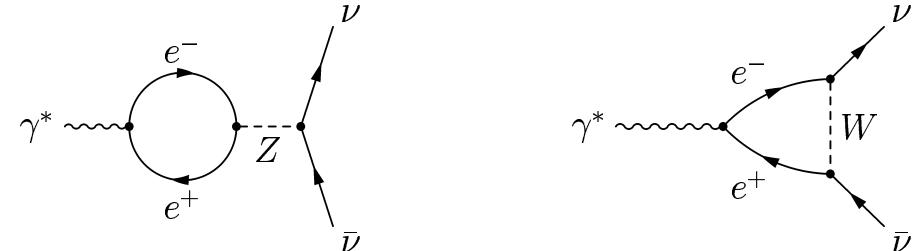
$$\frac{\partial f}{\partial t} + v^i \frac{\partial f}{\partial x^i} + \frac{\partial(fF^i)}{\partial p^i} = B_{EA}(f) + B_{NES}(f) + B_{\nu N}(f) + B_{TP}(f)$$
$$B(f) = \left[1 - f\right] \frac{1}{(2\pi)^3} \int_0^\infty E_2^2 dE_2 \int_{-1}^1 d\mu_2 \int_0^{2\pi} d\phi_2 R^p \left[1 - \bar{f}\right]$$
$$-f \frac{1}{(2\pi)^3} \int_0^\infty E_2^2 dE_2 \int_{-1}^1 d\mu_2 \int_0^{2\pi} d\phi_2 R^a \bar{f}$$

- Production/absorption kernels: $R^{p,a}(E_1, E_2, \cos \theta)$

$$R^p(E_1, E_2, \cos \theta) = \frac{8\pi^4}{E_1^2 E_2^2} \frac{d^3 \Gamma}{dE_1 dE_2 d \cos \theta}$$

The plasma neutrino process

- ▶ ‘massive’ photon
- ▶ 1-loop calculation



$$\mathcal{M} = \epsilon^\mu(K) \Gamma_{\mu\nu}(K) \left[\bar{u}_1(Q_1) \gamma^\nu (1 - \gamma_5) v_2(Q_2) \right]$$

$$\begin{aligned} |\mathcal{M}|^2 = & -V^{\mu\alpha} V_\mu{}^\beta v_{\alpha\beta} - A^{\mu\alpha} A_\mu{}^\beta v_{\alpha\beta} \\ & + (V^{\mu\beta} A_\mu{}^\alpha - V^{\mu\alpha} A_\mu{}^\beta) a_{\alpha\beta} \end{aligned}$$

- ▶ Vector (T+L), axial (T) and mixed (T) contributions

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Plasmon: $\langle |\mathcal{M}|^2 \rangle$

$$\langle |\mathcal{M}|^2 \rangle_T = \frac{G_F^2 (C_V^f)^2}{\pi \alpha} \Pi_T^2(\omega_T, k) \left[E_1 E_2 - \frac{(\mathbf{k} \cdot \mathbf{q}_1)(\mathbf{k} \cdot \mathbf{q}_2)}{k^2} \right],$$

$$\begin{aligned} \langle |\mathcal{M}|^2 \rangle_L &= 2 \frac{G_F^2 (C_V^f)^2}{\pi \alpha} \left(\frac{\omega_L^2 - k^2}{k^2} \right)^2 \Pi_L^2(\omega_L, k) \\ &\quad \times \left[\frac{(E_1 \omega_L - \mathbf{q}_1 \cdot \mathbf{k})(E_2 \omega_L - \mathbf{q}_2 \cdot \mathbf{k})}{\omega_L^2 - k^2} + \frac{(\mathbf{k} \cdot \mathbf{q}_1)(\mathbf{k} \cdot \mathbf{q}_2)}{k^2} \right. \\ &\quad \left. - \frac{E_1 E_2 + \mathbf{q}_1 \cdot \mathbf{q}_2}{2} \right], \end{aligned}$$

$$\langle |\mathcal{M}|^2 \rangle_A = \frac{G_F^2 (C_A^f)^2}{\pi \alpha} \Pi_A^2(\omega_T, k) \left[E_1 E_2 - \frac{(\mathbf{k} \cdot \mathbf{q}_1)(\mathbf{k} \cdot \mathbf{q}_2)}{k^2} \right],$$

$$\langle |\mathcal{M}|^2 \rangle_M = 2 \frac{G_F^2 C_A^f C_V^f}{\pi \alpha} \frac{\Pi_A(\omega_T, k) \Pi_T(\omega_T, k)}{k} \left[E_1(\mathbf{k} \cdot \mathbf{q}_2) - E_2(\mathbf{k} \cdot \mathbf{q}_1) \right].$$

Plasmon: production kernels (1)

- ▶ production and absorption kernels

$$\begin{aligned} R^p(E_1, E_2, \cos \theta) &= \frac{8\pi^4}{E_1^2 E_2^2} \frac{d^3 \Gamma}{dE_1 dE_2 d \cos \theta} \\ &= \frac{8\pi^4}{E_1^2 E_2^2 (E_1 + E_2)} \frac{d^3 Q}{dE_1 dE_2 d \cos \theta}. \end{aligned}$$

- ▶ Legendre expansion coefficients

$$R_a^p(E_1, E_2, \cos \theta) = \sum_{l=0}^{\infty} \frac{2l+1}{2} \Phi_l^a(E_1, E_2) P_l(\cos \theta),$$

Plasmon: production kernels (2)

$$\begin{aligned}
R_T^p(E_1, E_2, \cos \theta) &= \frac{G_F^2 (C_V^f)^2}{2\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_T^2(\omega_T, k)}{\omega_T} \\
&\quad \times \frac{(E_1 - E_2)^2 + E_1 E_2 (1 + \cos \theta)}{E_1^2 + E_2^2 + 2E_1 E_2 \cos \theta} (1 - \cos \theta) \delta(\omega_T - E_1 - E_2), \\
R_L^p(E_1, E_2, \cos \theta) &= \frac{G_F^2 (C_V^f)^2}{2\alpha} Z_L(k) n_B(\omega_L, T) (\omega_L^2 - k^2)^2 \frac{E_1 E_2}{\omega_L} \frac{1 - \cos^2 \theta}{k^2} \\
&\quad \times \delta(\omega_L - E_1 - E_2) \Theta(\mathcal{K}), \\
R_A^p(E_1, E_2, \cos \theta) &= \frac{G_F^2 (C_A^f)^2}{2\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_A^2(\omega_T, k)}{\omega_T} \\
&\quad \times \frac{(E_1 - E_2)^2 + E_1 E_2 (1 + \cos \theta)}{E_1^2 + E_2^2 + 2E_1 E_2 \cos \theta} (1 - \cos \theta) \delta(\omega_T - E_1 - E_2), \\
R_M^p(E_1, E_2, \cos \theta) &= \frac{G_F^2 C_A^f C_V^f}{\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_A(\omega_T, k) \Pi_T(\omega_T, k)}{\omega_T k} (E_2 - E_1) \\
&\quad \times (1 - \cos \theta) \delta(\omega_T - E_1 - E_2).
\end{aligned}$$

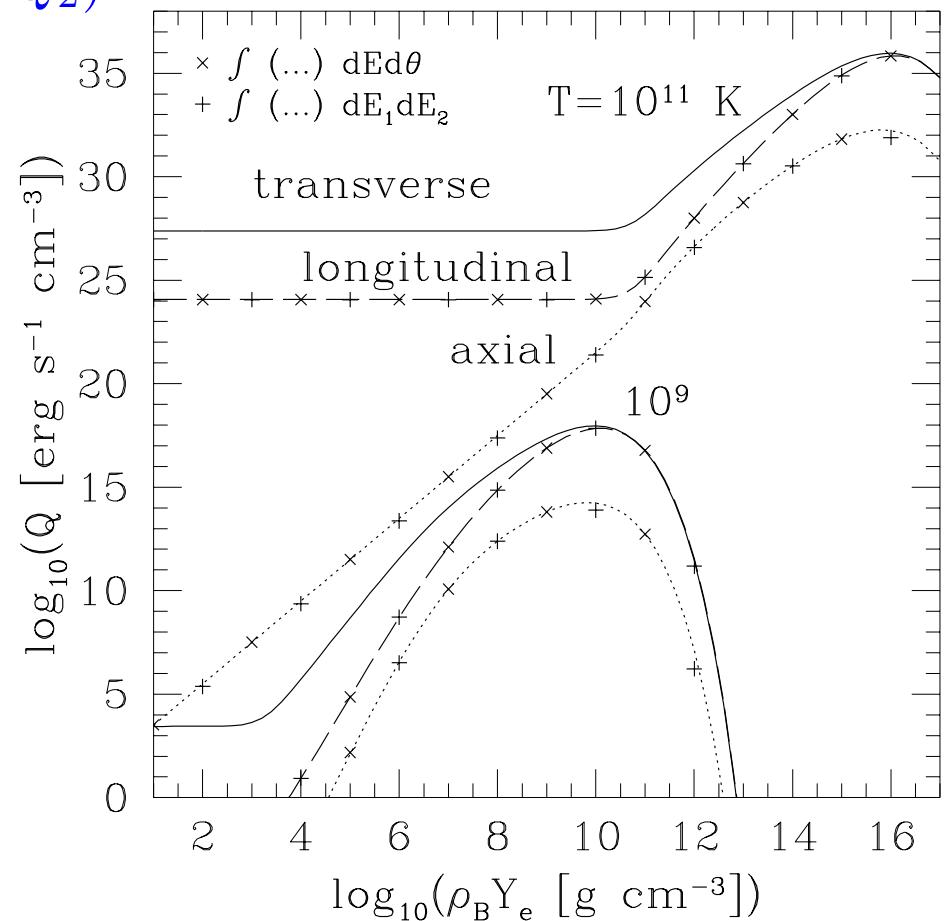
Plasmon: production kernels (3)

$$\begin{aligned}
\Phi_T^p(E_1, E_2) &= \frac{G_F^2 \sum_f (C_V^f)^2}{2\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_T^2(\omega_T, k)}{\omega_T E_1 E_2} \left[E_1 E_2 \right. \\
&\quad \left. - \frac{(k^2 + E_1^2 - E_2^2)(k^2 - E_1^2 + E_2^2)}{4k^2} \right] J_T(E_1, E_2) \Theta(4E_1 E_2 - \Pi_T) P_l(\cos \tilde{\theta}), \\
\Phi_L^p(E_1, E_2) &= \frac{G_F^2 \sum_f (C_V^f)^2}{2\alpha} Z_L(k) n_B(\omega_L, T) \frac{(\omega_L^2 - k^2)^2}{\omega_L} \frac{E_1 E_2 (1 - \cos^2 \tilde{\theta})}{E_1^2 + E_2^2 + 2E_1 E_2 \cos \tilde{\theta}} \\
&\quad \times J_L(E_1, E_2) \Theta(\mathcal{K}) \Theta(4E_1 E_2 - (\omega_L^2 - k^2)) P_l(\cos \tilde{\theta}), \\
\Phi_A^p(E_1, E_2) &= \frac{G_F^2 \sum_f (C_A^f)^2}{2\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_A^2(\omega_T, k)}{\omega_T E_1 E_2} \left[E_1 E_2 \right. \\
&\quad \left. - \frac{(k^2 + E_1^2 - E_2^2)(k^2 - E_1^2 + E_2^2)}{4k^2} \right] J_T(E_1, E_2) \Theta(4E_1 E_2 - \Pi_T) P_l(\cos \tilde{\theta}), \\
\Phi_M^p(E_1, E_2) &= \frac{G_F^2 \sum_f C_A^f C_V^f}{2\alpha} Z_T(k) n_B(\omega_T, T) \frac{\Pi_A(\omega_T, k) \Pi_T^2(\omega_T, k)}{k \omega_T E_1 E_2} (E_2 - E_1) \\
&\quad \times J_T(E_1, E_2) \Theta(4E_1 E_2 - \Pi_T) P_l(\cos \tilde{\theta})
\end{aligned}$$

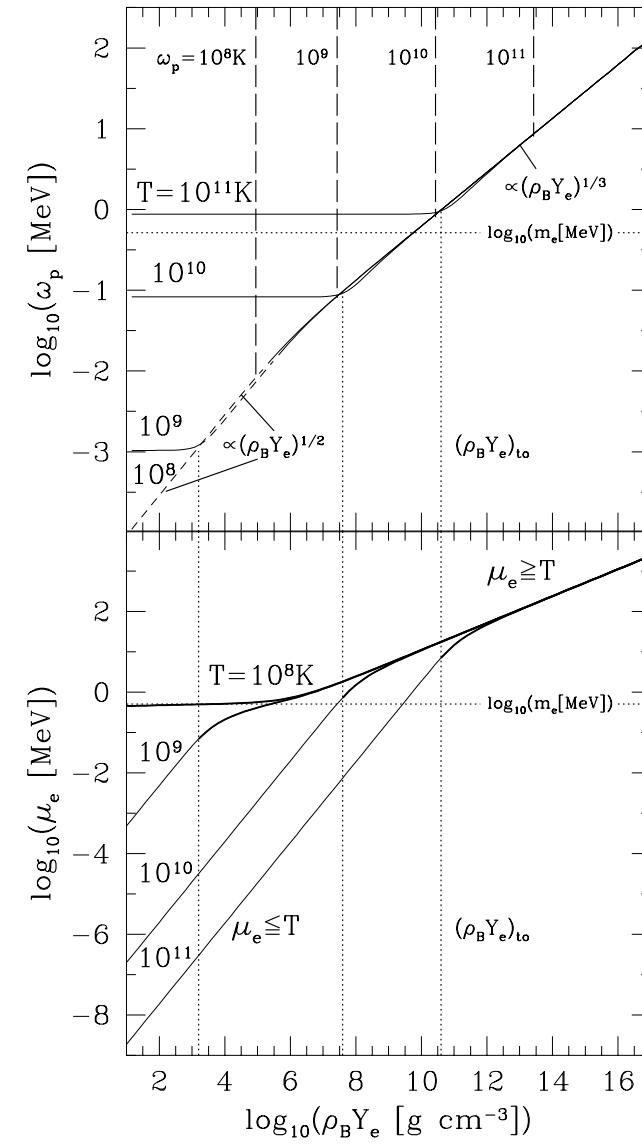
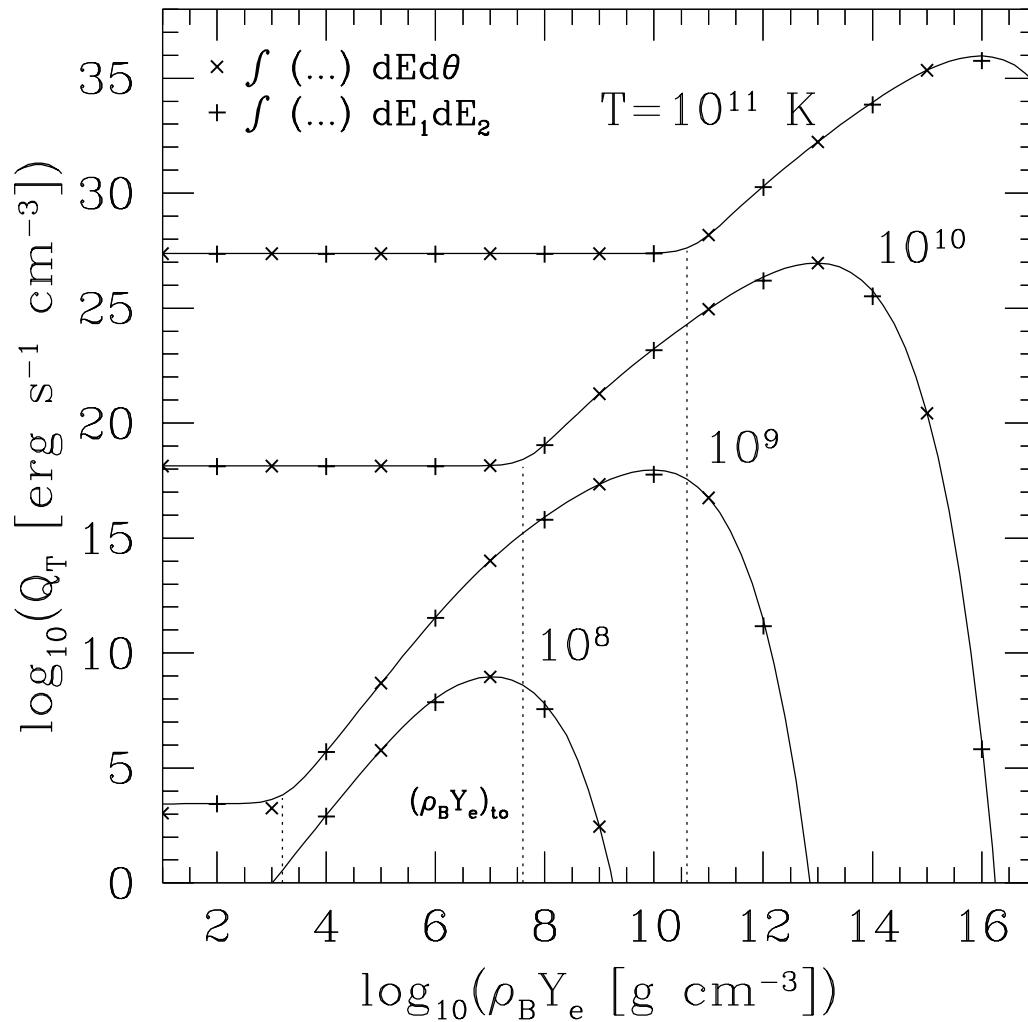
Plasmon: Total emissivity

$$Q = \sum_{\epsilon} \int \frac{d^3 k}{2\omega(2\pi)^3} Z(k) \frac{d^3 q_1}{2E_1(2\pi)^3} \frac{d^3 q_2}{2E_2(2\pi)^3} (E_1 + E_2) \langle |\mathcal{M}|^2 \rangle \\ \times n_B(\omega, T) (2\pi)^4 \delta^4(K - Q_1 - Q_2)$$

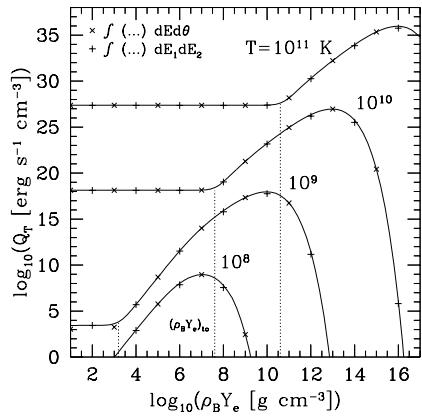
- ▶ Differential Q 's integrated
- ▶ Q_T dominates
- ▶ $Q_L \simeq Q_T$ at high ρ
- ▶ Q_A negligible



Plasmon: T, ρ dependence

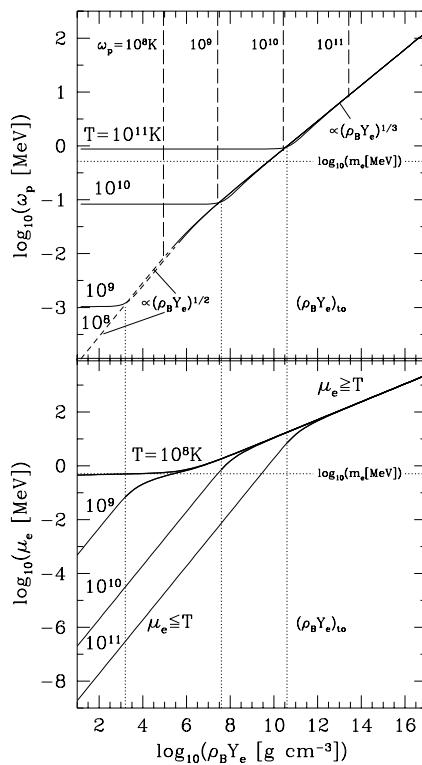


High temperature regime ($T \gg \omega_p$)



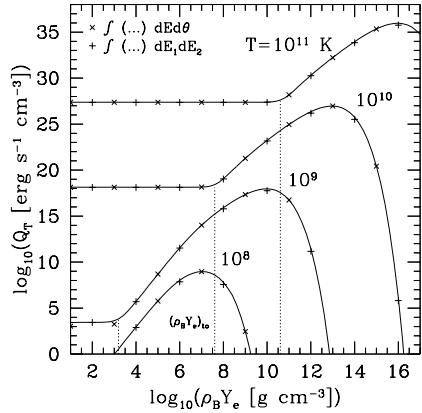
► emissivity

$$Q_T \sim 4 \frac{\sum_f C_V^f G_F}{96\pi^4 \alpha} \zeta(3) \omega_p^6 T^3$$



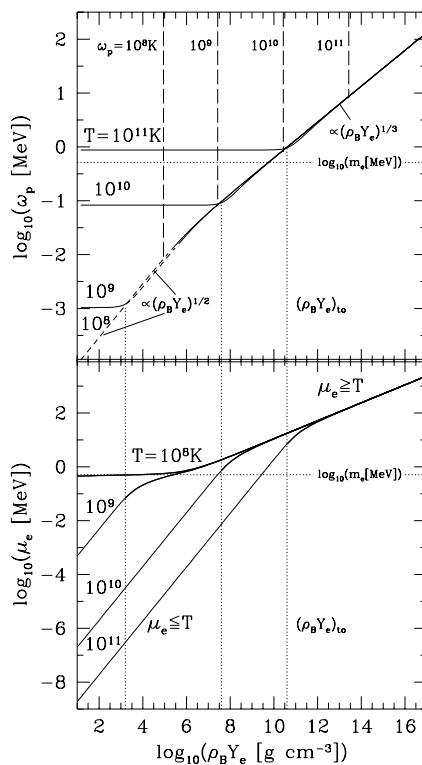
- plasma frequency $\omega_p^2 \approx \frac{4\alpha\pi}{9} T^2$
- ω_p and Q_T effectively n_e -independent

High density regime ($T \ll \omega_p$)



- emissivity:

$$Q_T \sim 2 \frac{\sum_f C_V^f G_F}{96\pi^4 \alpha} \sqrt{\frac{\pi}{2}} \omega_p^{15/2} T^{3/2} e^{-\omega_p/T}$$

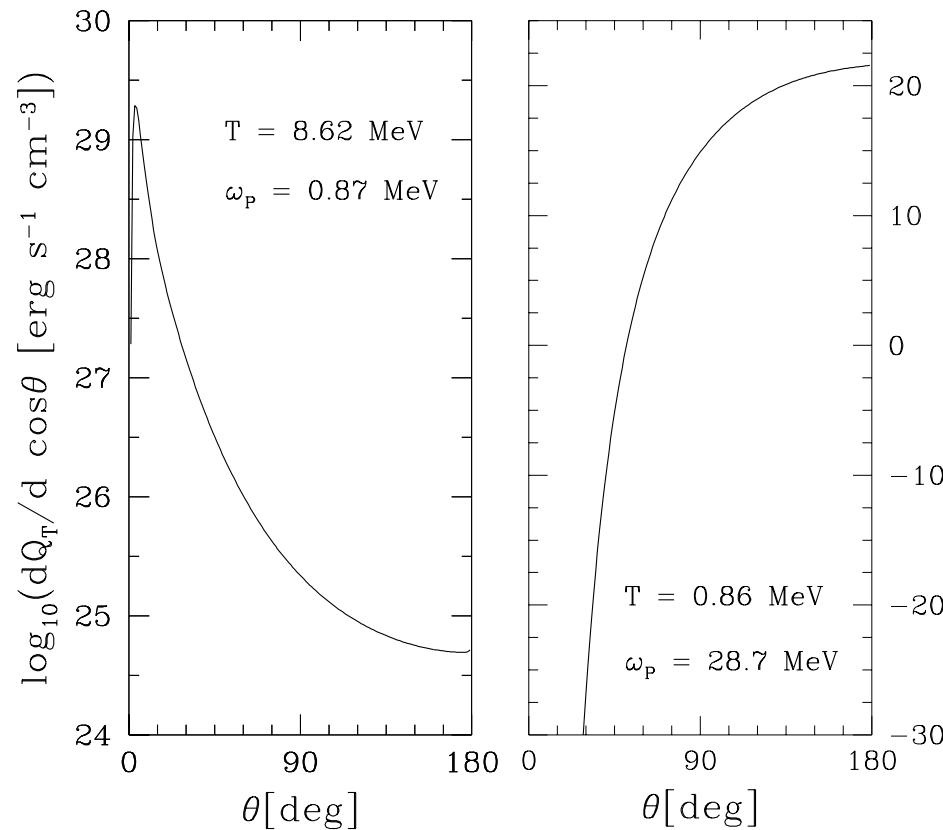


- ω_p depends strongly on n_e

classical: $\omega_p \sim n_e^{1/2}$ degenerate: $\omega_p \sim n_e^{1/3}$

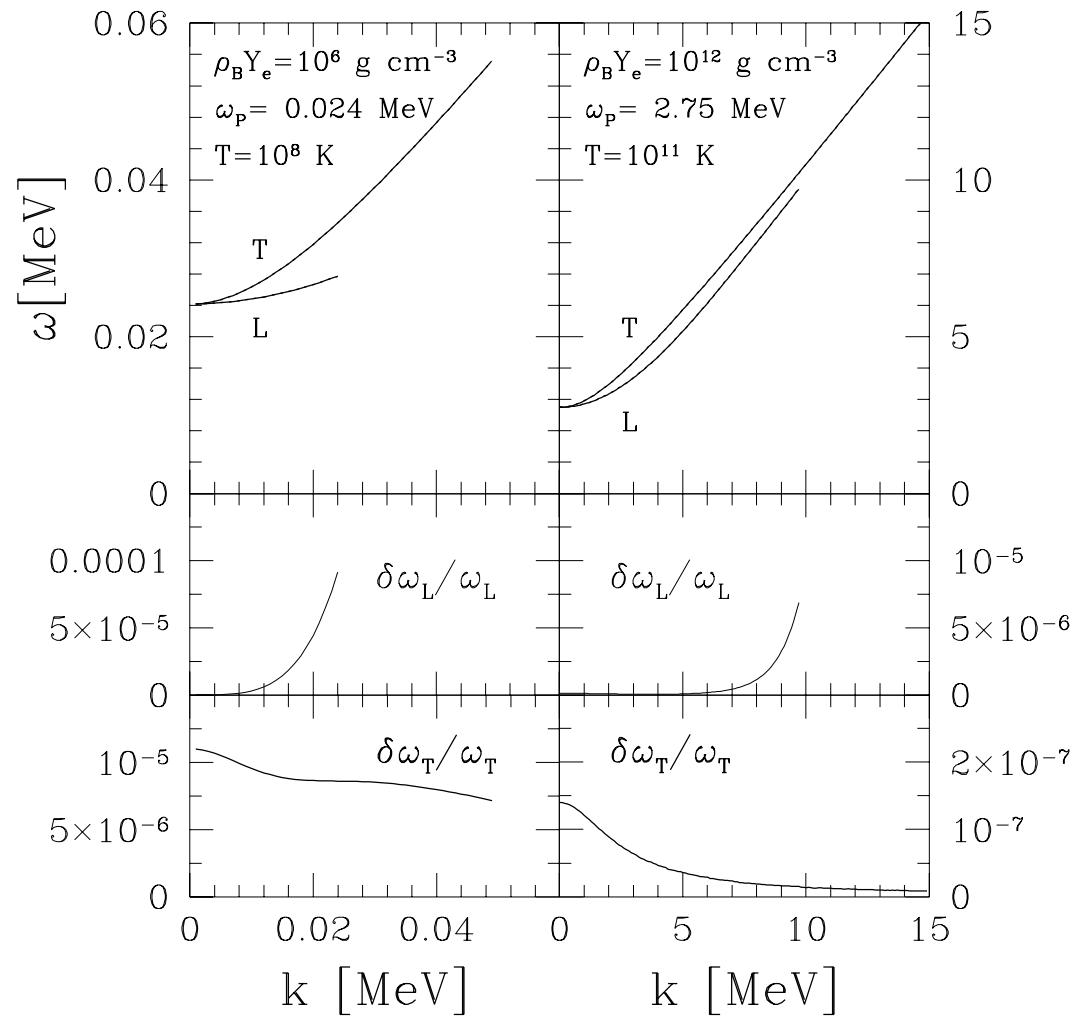
- strong density dependence of Q
- rise of emissivity with net electron density
- exponential damping $e^{-\omega_p/T}$

Regimes (cont.)



- ▶ $\omega_p \gg T$ – massive particle \Rightarrow “back to back”
- ▶ $\omega_p \ll T$ – light particle \Rightarrow small angle

Approximate vs. exact



Plasma neutrino: Summary

- ▶ full $|\mathcal{M}|^2$ is computed
- ▶ $|\mathcal{M}|^2$
 - transverse (**vector**, axial, mixed)
 - longitudinal (vector only) – high ρ
- ▶ BTE relevant quantities:
 $R^{p,a}(E_1, E_2, \theta), \Phi_l^{p,a}(E_1, E_2)$
- ▶ 1-loop dispersion relation
 - accurate approximation
(Braaten, Segel: PRD47, 1993)
- ▶ analysis of the process in different regimes

The photo-neutrino process

► $\gamma^* + e^\pm \rightarrow e^\pm + \nu + \bar{\nu}$

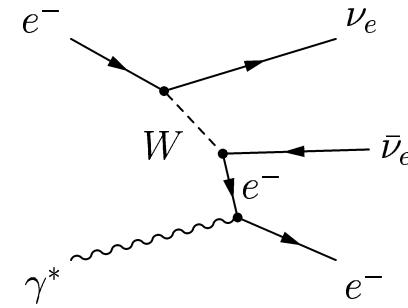
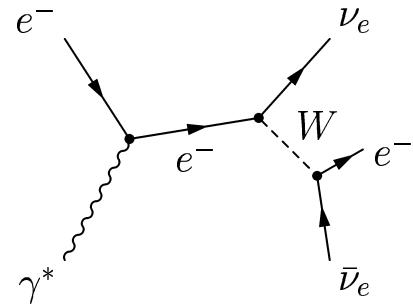
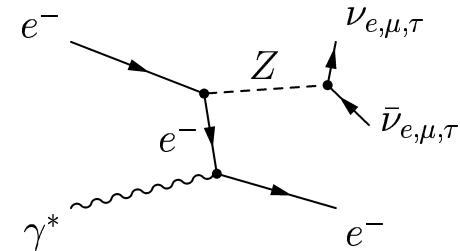
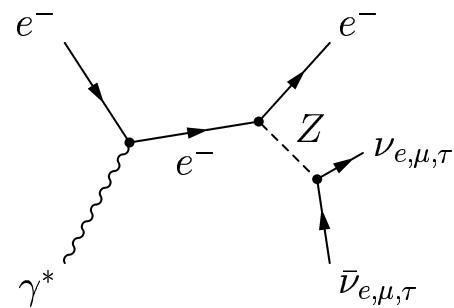


Photo-neutrino

- ▶ full $|\mathcal{M}|^2$
- ▶ dQ 's, $d\Gamma$'s
- ▶ $R^{p,a}$, $\Phi^{p,a}$
- ▶ verified through total Q's
- ▶ lowest order dispersion relation
- ▶ accurate for nonrelativistic & nondegenerate matter

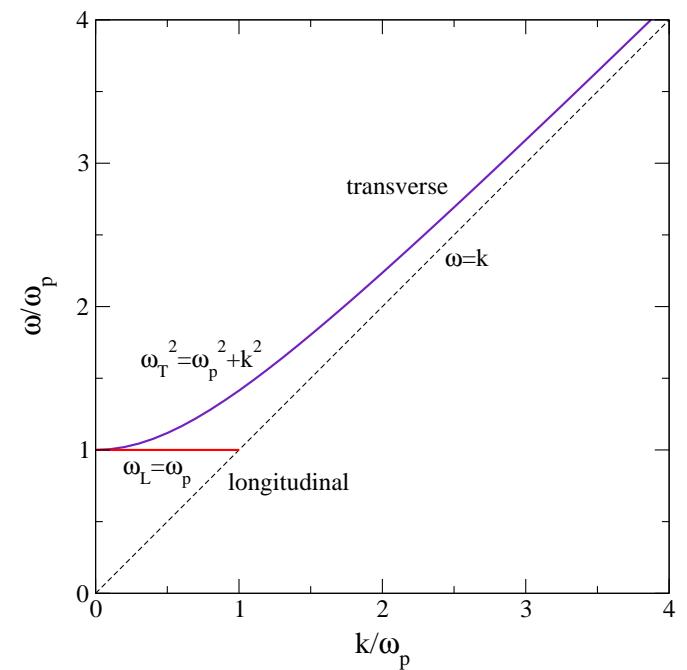


Photo-neutrino: $R^{p,a}(E_1, E_2, \theta)$,

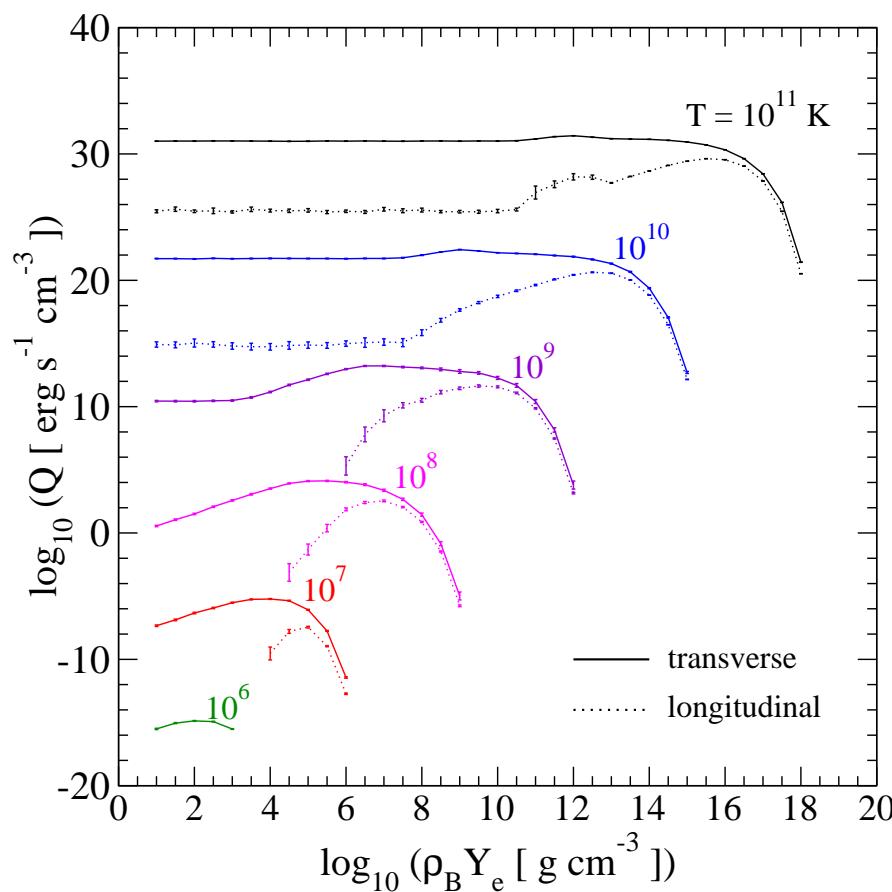
$\Phi_l^{p,a}(E_1, E_2)$

$$R^p(E_q, E_{q'}, \cos \theta_{qq'}) = \int \frac{2d^3\mathbf{p}}{(2\pi)^3} \frac{F_e(E_p)}{2E_p} \int \frac{2d^3\mathbf{k}}{(2\pi)^3} \frac{F_\gamma(\omega)}{2\omega} \int \frac{d^3\mathbf{p}'}{(2\pi)^3} \frac{[1 - F_e(E_{p'})]}{2E_{p'}} \\ \times \frac{(2\pi)^4}{4E_q E_{q'}} \delta^4(p + k - p' - q - q') \sum_{s,=e} |\mathcal{M}|^2$$

$$\Phi_l^{p,a}(E_q, E_{q'}) = \int_{-1}^1 d(\cos \theta_{qq'}) P_l(\cos \theta_{qq'}) R^p_a(E_q, E_{q'}, \cos \theta_{qq'}) .$$

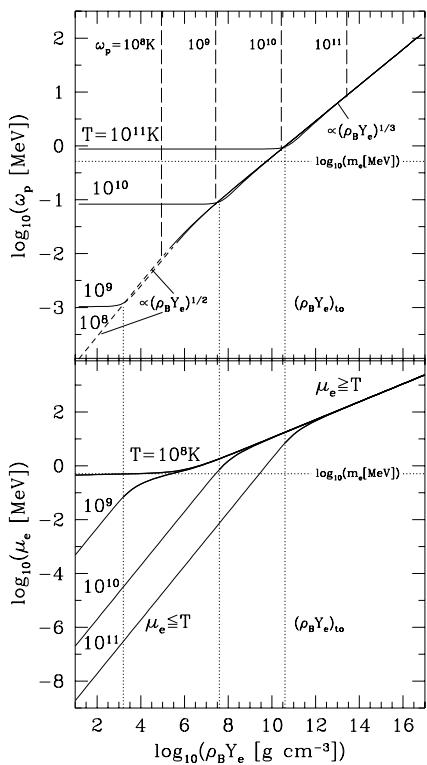
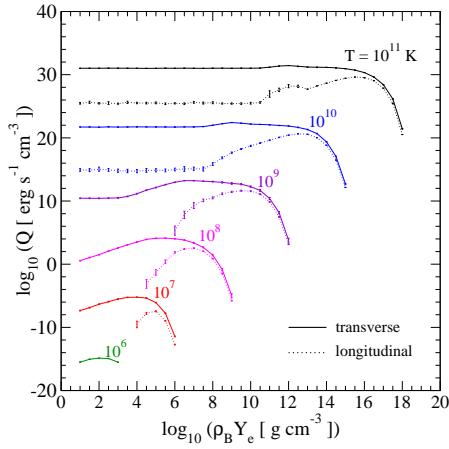
Photo-neutrino: Total emissivity

$$Q = \sum_{\epsilon} \int \frac{d^3 k}{2\omega(2\pi)^3} Z(k) \frac{d^3 q_1}{2E_1(2\pi)^3} \frac{d^3 q_2}{2E_2(2\pi)^3} \frac{d^3 p_1}{2E_{e1}(2\pi)^3} \frac{d^3 p_2}{2E_{e2}(2\pi)^3} \\ \times (E_1 + E_2) \langle |\mathcal{M}|^2 \rangle n_B n_{F1} (1 - n_{F2}) (2\pi)^4 \delta^4(P_1 + K - P_2 - Q_1 - Q_2)$$



(In preparation)

Degenerate limit

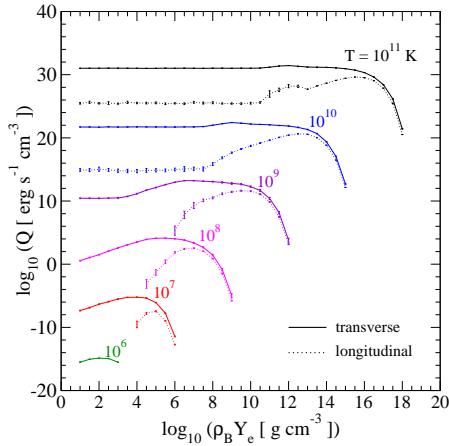


- ▶ $\mu_e \gg T, \mu_e \gg m_e, \omega_P \gg T, \omega_P \gg m_e$
- ▶ energy scales
 - $E_p \simeq E_{p'} \simeq |\mathbf{p}| \simeq |\mathbf{p}'| \simeq \mu_e$
 - $\omega \simeq \omega_p \simeq E_q + E_{q'}, \quad |\mathbf{k}| \simeq T$
 - $E_q \simeq |\mathbf{q}| \simeq \omega_p/2$
- ▶ emissivity

$$Q_T \simeq \frac{4}{3} \frac{\alpha G_F^2 (C_V^2 + C_A^2)}{(2\pi)^6} \omega_p^6 T^3 e^{-\omega_p/T}$$

$$(\rho_B Y_e)_{max} \simeq 9.213 \times 10^{12} \left(\frac{T}{\text{MeV}} \right)^3 \text{ g cm}^{-3}$$

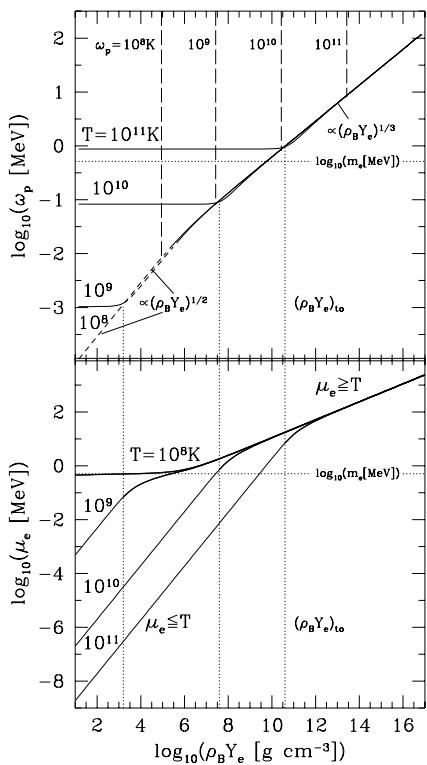
Nondegenerate limit



- energy scales $\sim T$

$$Q_T \simeq 4.4 \times 10^{22} \left(\frac{T}{\text{MeV}} \right)^9 \text{erg s}^{-1} \text{cm}^{-3}$$

$$\log_{10}(Q_T) = \begin{cases} 31.05 & \text{at } T = 8.6 \text{ MeV (}10^{11} \text{ K)} \\ 22.05 & \text{at } T = 0.86 \text{ MeV} \end{cases}$$



- turn-on at $\mu_e \simeq T$

$$(n_e)_{\text{to}} \simeq \frac{T^3}{2.72} \quad \text{or}$$

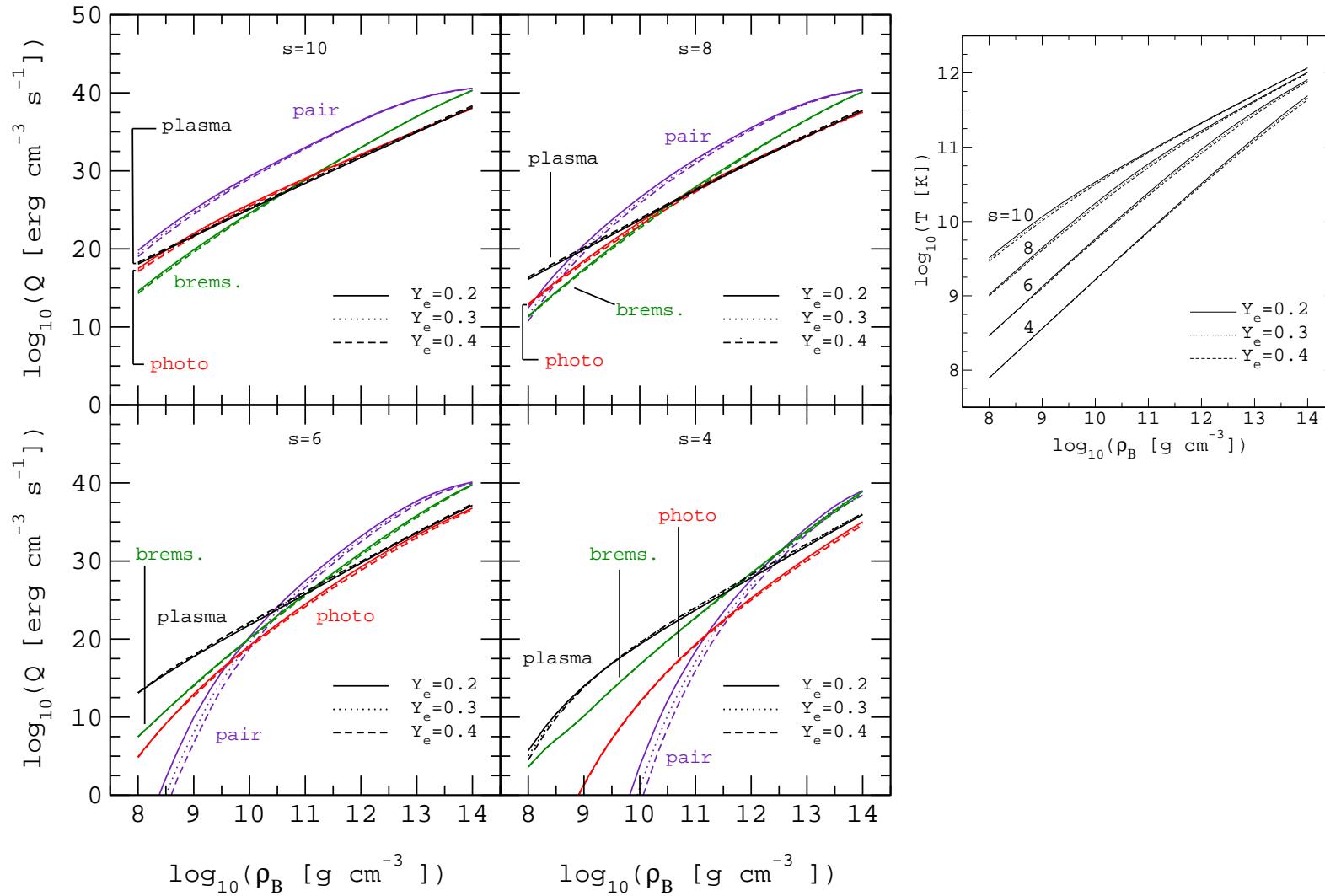
$$(\rho_B Y_e)_{\text{to}} \simeq 8 \times 10^7 \left(\frac{T}{\text{MeV}} \right)^3 \text{g cm}^{-3}$$

$$\Rightarrow \log_{10}(\rho_B Y_e) = \begin{cases} 10.7 & \text{at } T = 8.6 \text{ MeV (}10^{11} \text{ K)} \\ 7.7 & \text{at } T = 0.86 \text{ MeV} \end{cases}$$

Photo-neutrino: Summary

- ▶ full $|\mathcal{M}|^2$ is computed
- ▶ $|\mathcal{M}|^2$
 - transverse
 - longitudinal - high ρ
 - verified: Q, Γ
 - $R^{p,a}(E1, E2, \theta), \Phi_l^{p,a}(E1, E2)$
- ▶ approximations for different physical regimes

Comparison with competing processes



Summary

- ▶ computed full $|\mathcal{M}|^2$'s
- ▶ analytical expressions for dQ and $d\Gamma$'s
- ▶ production/absorption kernels for BTE
- ▶ relative importance depends on T , ρ
- ▶ importance to be explored
(neutrino transport + hydrodynamics)

